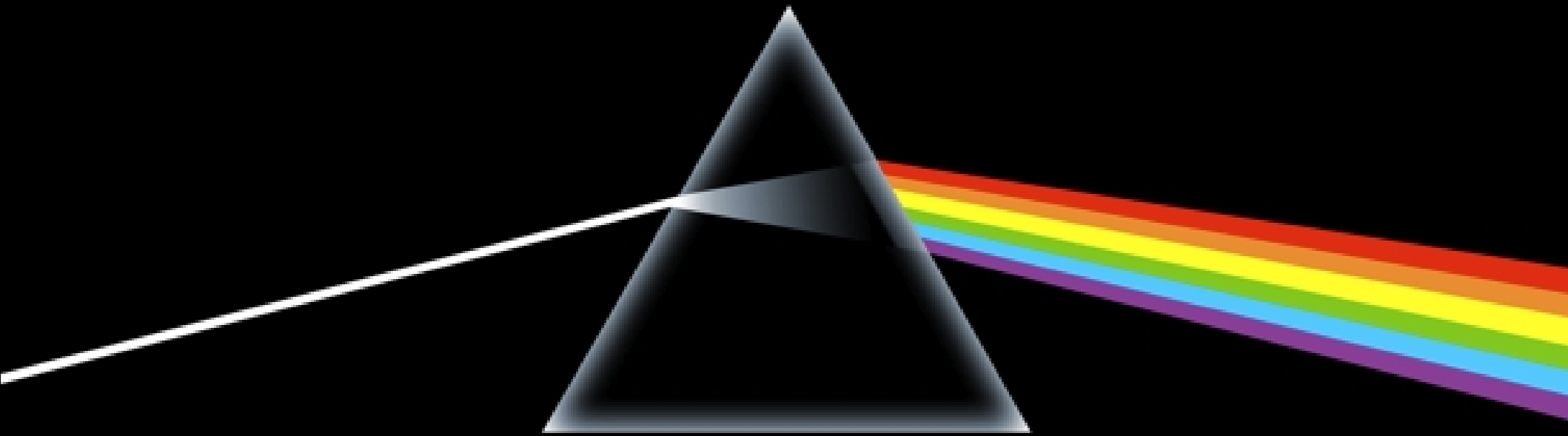


# Spectra of Gases and Solids



# Why are we doing this lab?

Spectra are unique to each element. This allows you to ID an element based on the spectral lines you see (like a fingerprint).

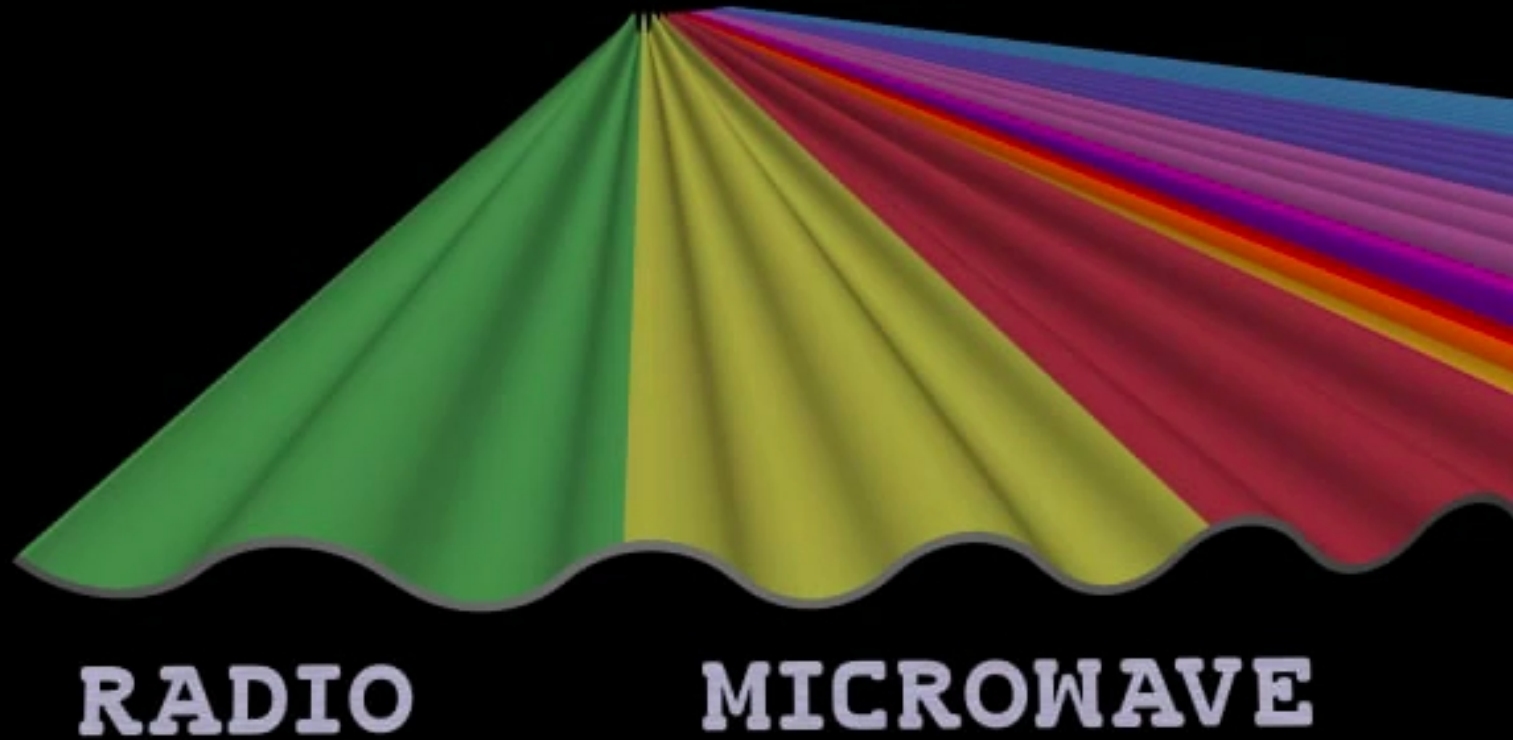
This is particularly useful in astronomy since we can't go to a star to figure out what it's made of. Just from the light we receive, we can determine a star's composition.



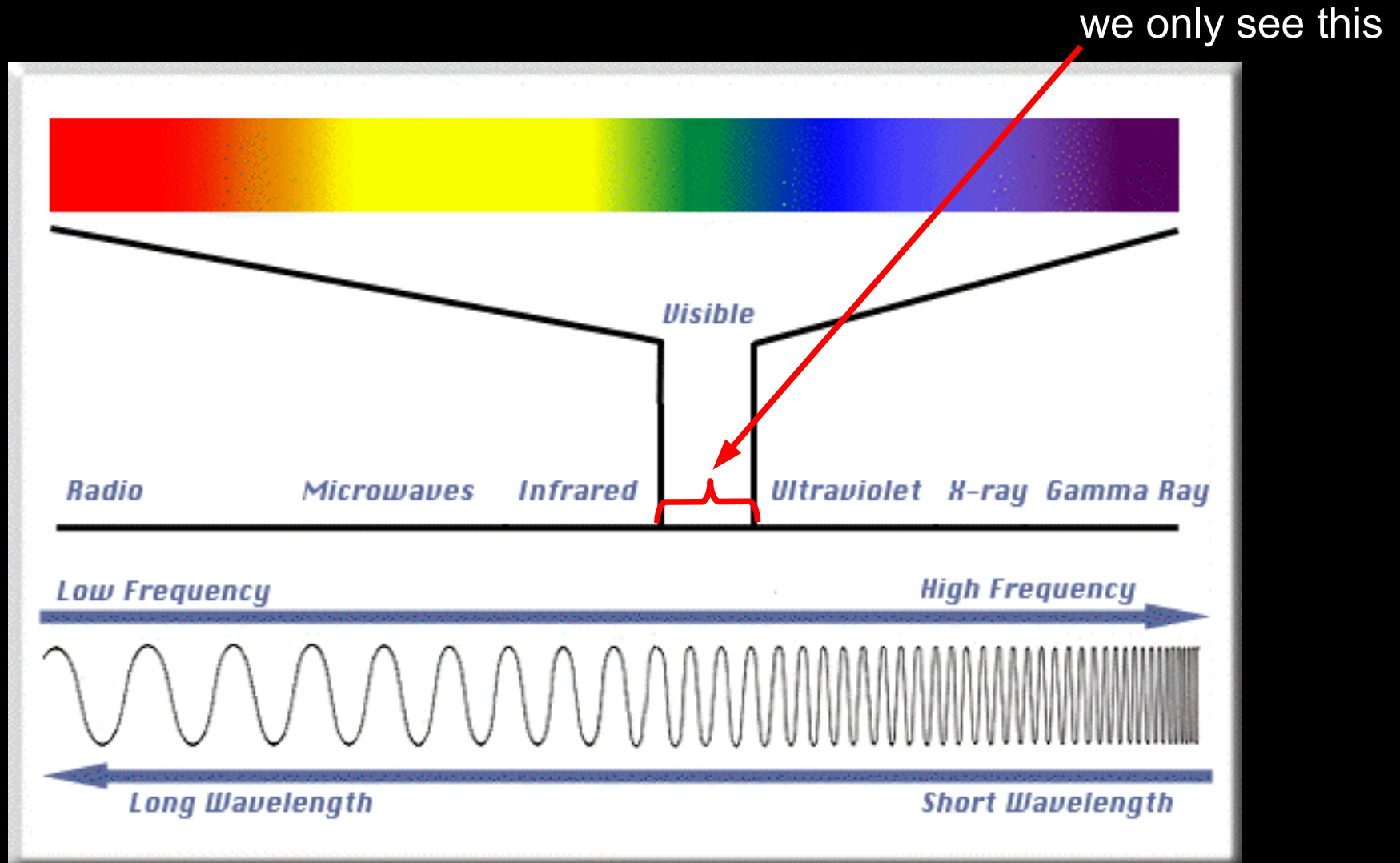
# What is a Spectrum?



# What is light?



# What is light?



The speed of light,  $c$ , is  $3.00 \times 10^8$  m/s (a constant). And  $c = \lambda f$ .  
Thus, longer wavelengths have smaller frequencies (and vice versa).

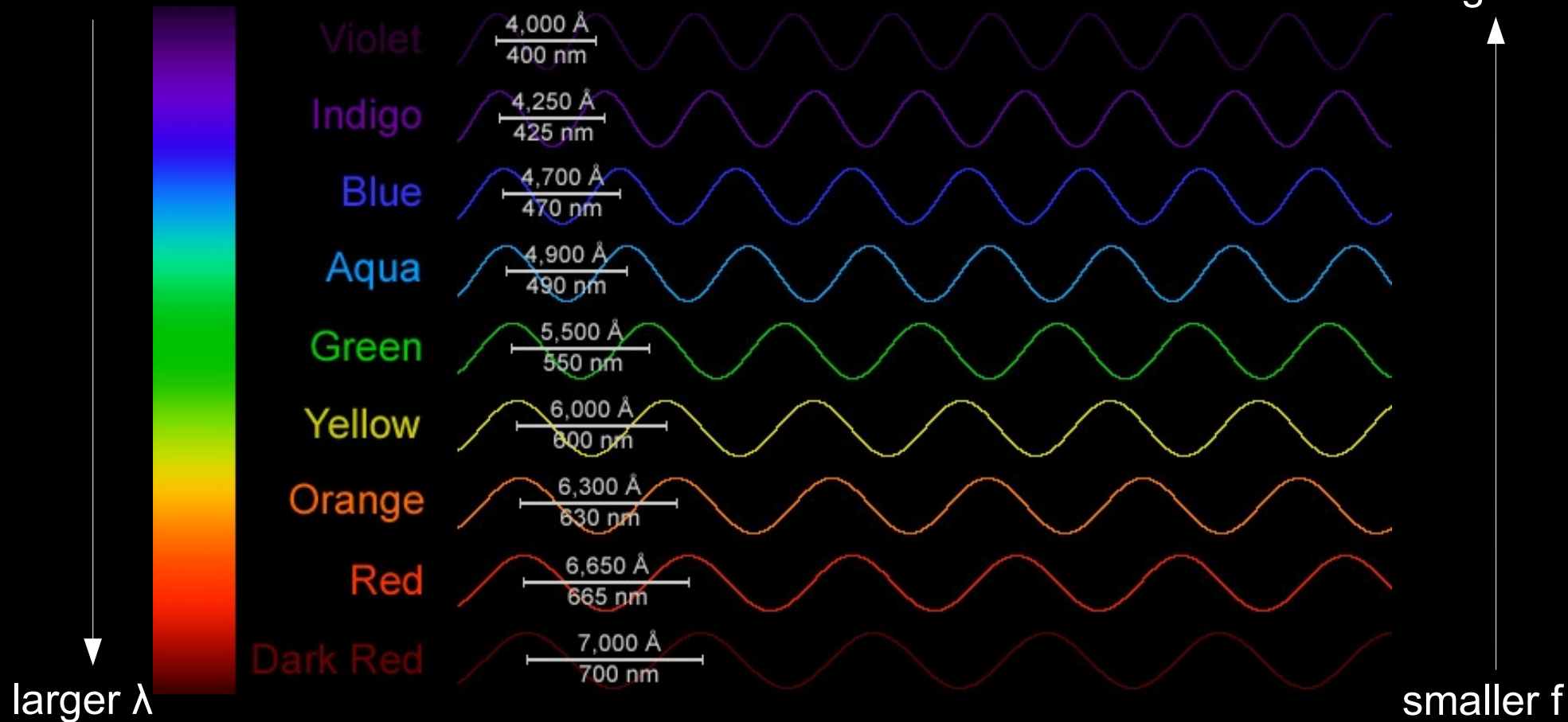


# Visible Light

Red light is lower frequency (lower energy) than blue light.

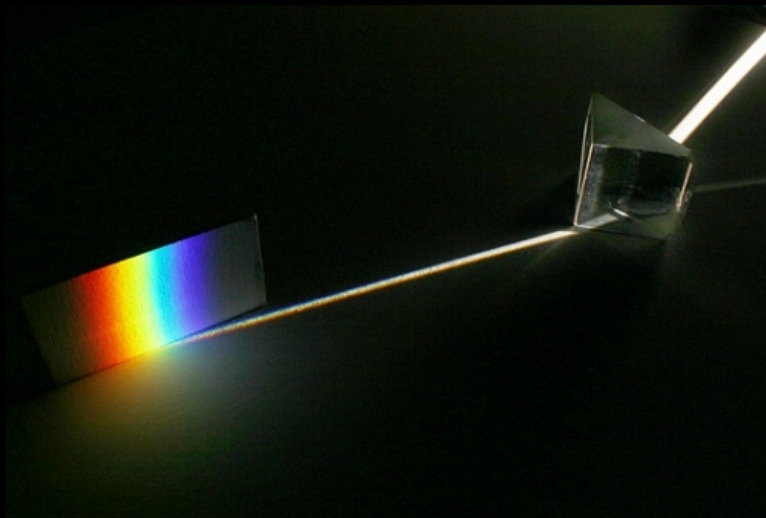
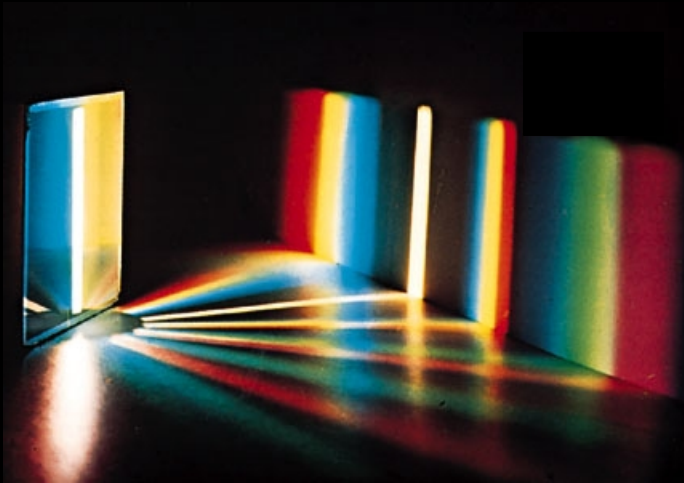
smaller  $\lambda$

larger  $f$



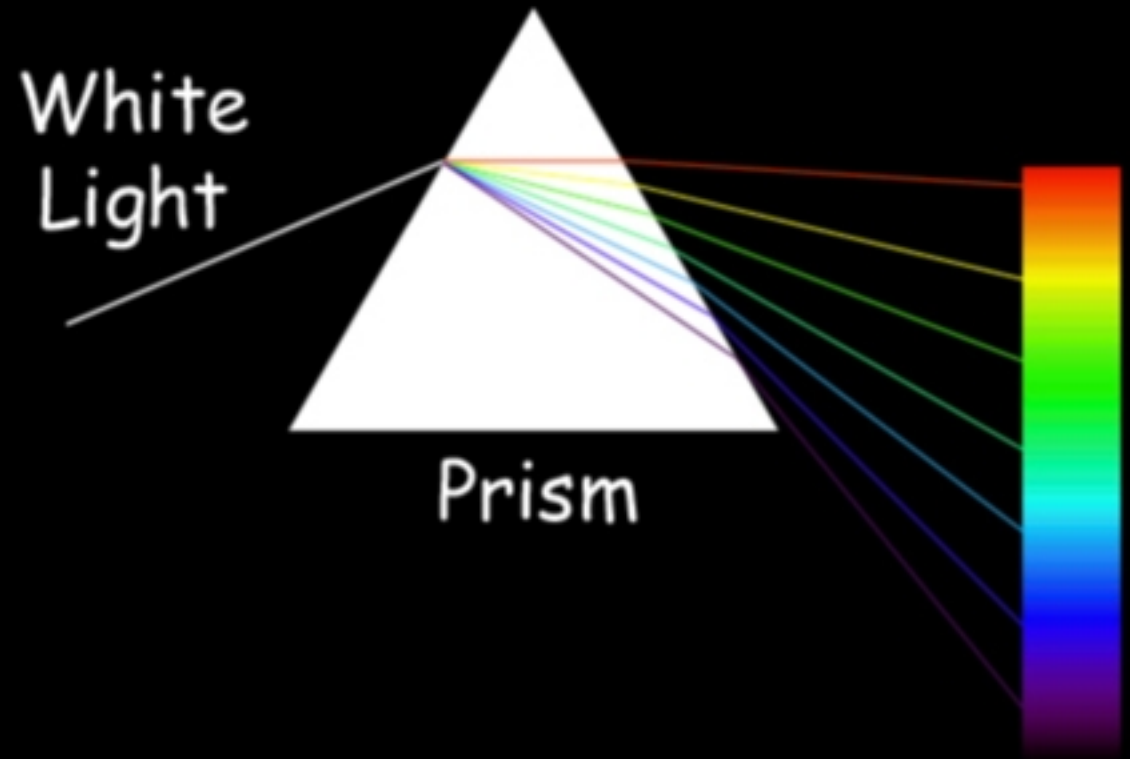
# How do you make a Spectrum?

If you bend light (spread it) different wavelengths will bend at different angles. That lets you see a “rainbow”.



# Continuous Spectra

Has all the colours of a rainbow...

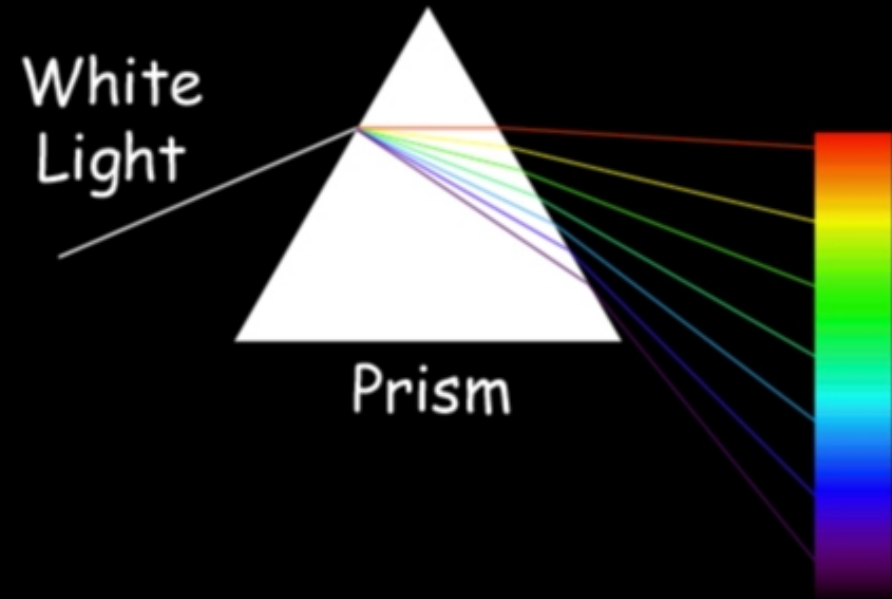
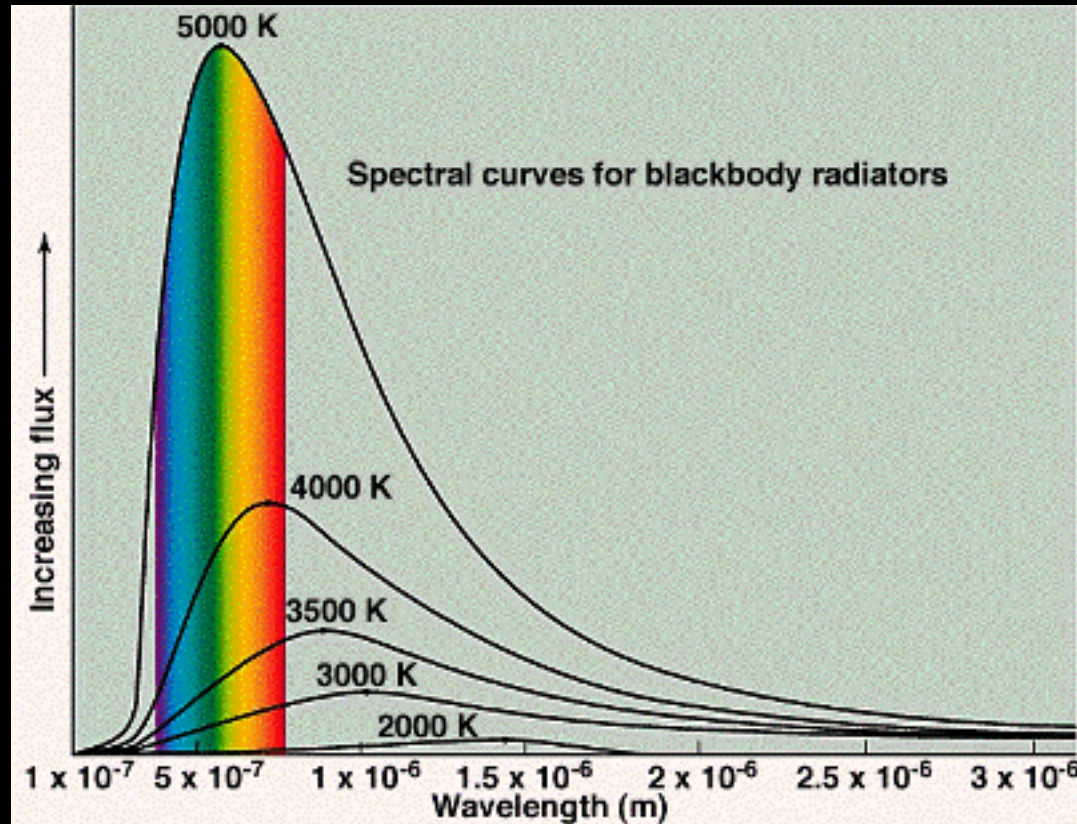


If you plot the brightness at each wavelength, you create a special curve.



# Continuous Spectra

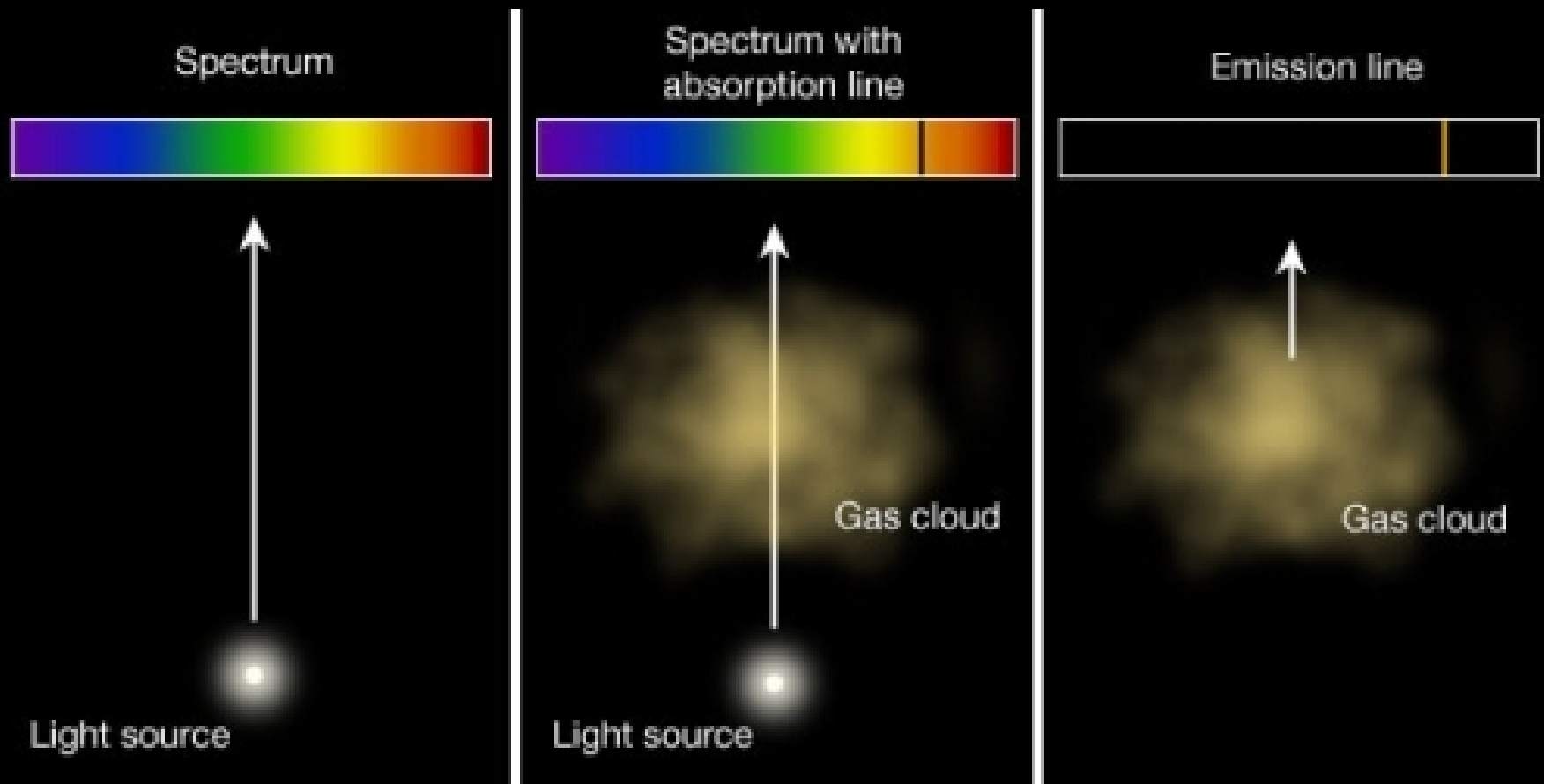
The special curve...



The amount of light and type of light you see depends on the source's temperature!

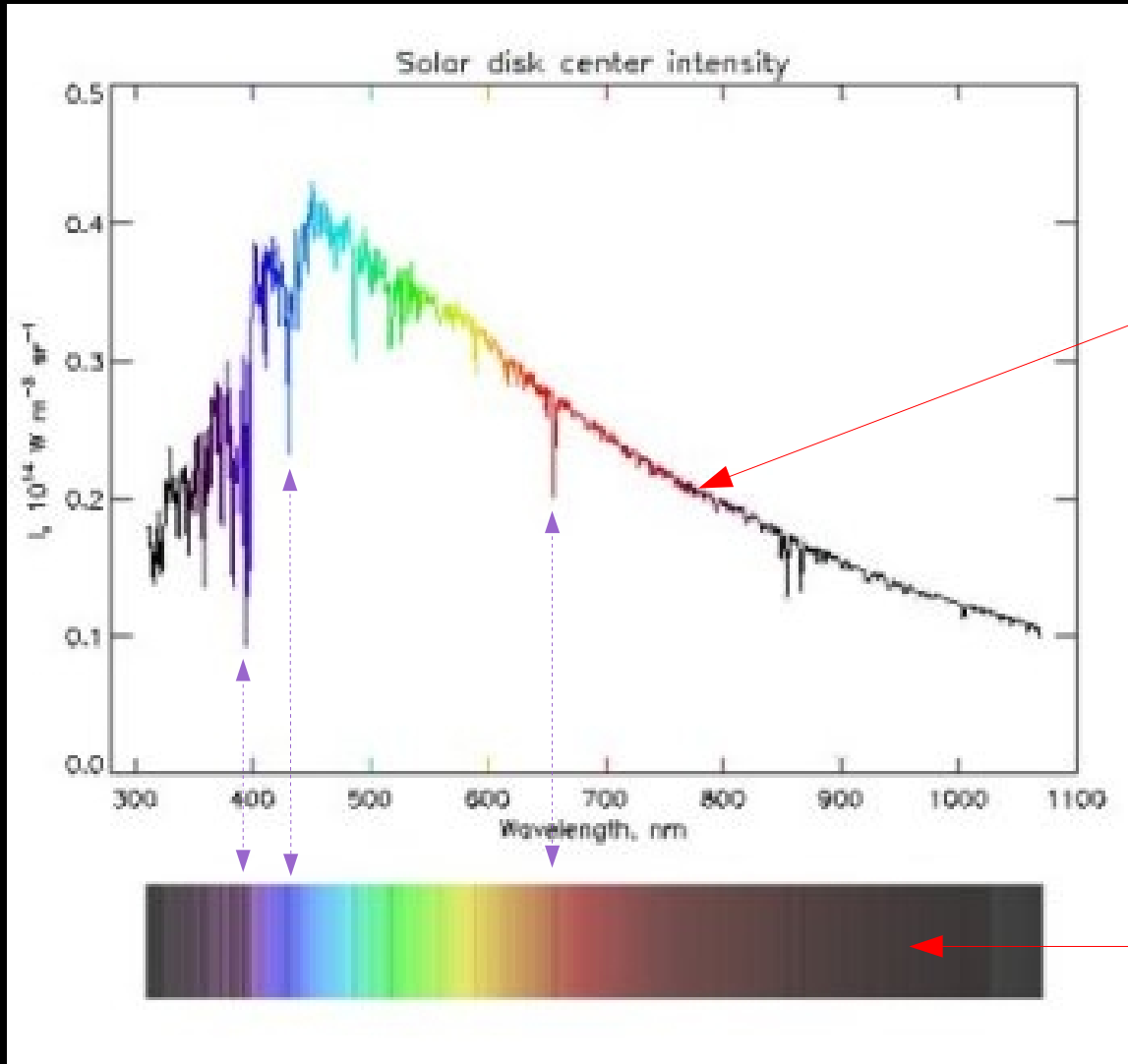
# Kirchoff's Laws

Depending on how the light gets to us, we see a different fingerprint pattern.



# The Sun

## The Sun's spectrum



The Sun's “black body” curve. Note that some colours are removed.

Sun's spectrum  
Darker absorption lines mean less light reached us.

